

Real Gas Computation Using an Energy Relaxation Method and High-Order WENO Schemes

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In this paper, we use a recently developed energy relaxation theory by Coquel and Perthame and high order weighted essentially nonoscillatory (WENO) schemes to simulate the Euler equations of real gas. The main idea is an energy decomposition under the form $\varepsilon = \varepsilon_1 + \varepsilon_2$, where ε_1 is associated with a simpler pressure law (γ -law in this paper) and the nonlinear deviation ε_2 is convected with the flow. A relaxation process is performed for each time step to ensure that the original pressure law is satisfied. The necessary characteristic decomposition for the high order WENO schemes is performed on the characteristic fields based on the ε_1 γ -law. The algorithm only calls for the original pressure law once per grid point per time step, without the need to compute its derivatives or any Riemann solvers. Both one- and two-dimensional numerical examples are shown to illustrate the effectiveness of this approach. © 1999 Academic Press

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1. INTRODUCTION

In this paper we consider the Euler equations for a real compressible inviscid fluid,

$$\partial_t \rho + \operatorname{div}(\rho u) = 0, \quad t \geq 0, x \in \mathbb{R}^d,$$

$$\partial_t \rho u + \operatorname{div}(\rho u \otimes u + p) = 0,$$

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$$\begin{aligned}\partial_t E + \operatorname{div}((E + p)u) &= 0, \\ E &= \frac{1}{2}\rho|u|^2 + \rho\varepsilon,\end{aligned}\tag{1.1}$$

where the quantities ρ , u , p , E , and ε represent the density, velocity, pressure, total energy, and specific internal energy, respectively. In addition, there is an equation of state (EOS) of the form $p = p(\rho, \varepsilon)$ associated with a strictly convex entropy $\rho s(\rho, \varepsilon)$ which satisfies the entropy inequalities

$$\partial_t \rho s + \operatorname{div}(\rho s u) \leq 0.\tag{1.2}$$

The pressure law is, furthermore, assumed to satisfy

$$p_{,\varepsilon}(\rho, \varepsilon) > 0, \quad p(\rho, 0) = 0, \quad p(\rho, \infty) = \infty.\tag{1.3}$$

In the literature research has been done in order to extend classical schemes designed for perfect gas to real gases. Collela and Glaz [1] extended the numerical procedure for obtaining the exact Riemann solution to a real-gas case, Grossman and Walters [7], Liou, van Leer, and Shuen [13] extended the method of flux-vector splitting and flux-difference splitting, Montagné, Yee, and Vinokur [16] developed second-order explicit shock-capturing schemes for real gas, Glaister [5] presented an extension of the approximate linearized Riemann solver with different averaged matrices, while Loh and Liou [15] used the generalization of their Lagrangian approach (originally proposed for perfect gas) to obtain the real gas Riemann solution.

Most of the previous proposed methods would require a computation of the pressure law and its derivatives, or a Riemann solver. This is not only costly but also problematic when there are no analytical expressions of the pressure law (for example, if we have only table values).

Recently Coquel and Perthame [2] have introduced an energy relaxation theory for Euler equations of real gas. The main idea is to introduce a relaxation of the nonlinear pressure law by considering an energy decomposition under the form $\varepsilon = \varepsilon_1 + \varepsilon_2$. The internal energy ε_1 is associated with a simpler pressure law p_1 (which is taken as the γ -law in this paper), while ε_2 stands for the nonlinear perturbation and is simply convected by the flow. These two energies are also subject to a relaxation process, and in the limit of an infinite relaxation rate, one recovers the initial pressure law p .

From this general framework, Coquel and Perthame have also deduced the extension to general pressure laws of classical schemes for polytropic gases, which only uses a single call to the pressure law per grid point and time step. No derivatives of the pressure law or any Riemann solvers need to be computed. Another advantage of their approach is that its implementation does not depend on the particular expression of the equation of states. For the first-order Godunov scheme, they have shown that this extension satisfies stability, entropy, and accuracy conditions. Numerical examples have been provided using first-order schemes by In [9].

For Euler equations of polytropic gas, high order ENO (essentially nonoscillatory) and WENO (weighted ENO) schemes [8, 18–20, 14, 10, 21] have been quite successful in providing high resolution results for complicated flow structures and shocks. The idea of ENO schemes is to adaptively choose local stencils so that interpolation across a discontinuity is avoided as much as possible. WENO uses a nonlinear combination of all stencils to improve

upon accuracy and smoothness of numerical fluxes while maintaining the nonoscillatory behavior of ENO near a discontinuity.

The aim of this paper is to study the implementation of this relaxation method with high-order WENO schemes [10] for real gases. One- and two-dimensional numerical examples will be given.

In Section 2 we provide the general framework of the energy relaxation theory of [2], followed by a short description of high-order WENO schemes [10]. We then give the details of the construction of the relaxed WENO schemes for general gases. In Section 3 numerical examples are given. We start with a description of the different equations of states used in this paper, followed by one-dimensional shock tube test problems. Two-dimensional test cases of a smooth vortex, to test the accuracy of the schemes, and of the double Mach reflection problem are then presented. Concluding remarks are given in Section 4. In the appendices, we give the expressions of the Roe matrices for the relaxation system (A) and the two-molecular vibrating gas (B).

2. IMPLEMENTATION OF THE ENERGY RELAXATION METHOD WITH WENO

2.1. Energy Relaxation Theory

The principle of the energy relaxation theory developed by Coquel and Perthame [2] is to find a pressure law $p_1(\rho^\lambda, \varepsilon_1^\lambda)$ (simpler than p , typically a polytropic law) and an internal energy $\phi(\rho^\lambda, \varepsilon_1^\lambda)$ so that the system (1.1) and the entropy inequality (1.2) can be recovered, in the limit of an infinite relaxation rate λ (called the *equilibrium limit*), from the system (called the *relaxation system*)

$$\begin{aligned} \partial_t \rho^\lambda + \operatorname{div}(\rho^\lambda u^\lambda) &= 0, \quad t \geq 0, x \in \mathbb{R}^d, \\ \partial_t \rho^\lambda u^\lambda + \operatorname{div}(\rho^\lambda u^\lambda \otimes u^\lambda + p_1^\lambda) &= 0, \\ \partial_t E_1^\lambda + \operatorname{div}((E_1^\lambda + p_1^\lambda)u^\lambda) &= \lambda \rho^\lambda (\varepsilon_2^\lambda - \phi(\rho^\lambda, \varepsilon_1^\lambda)), \\ \partial_t \rho^\lambda \varepsilon_2^\lambda + \operatorname{div}(\rho^\lambda u^\lambda \varepsilon_2^\lambda) &= -\lambda \rho^\lambda (\varepsilon_2^\lambda - \phi(\rho^\lambda, \varepsilon_1^\lambda)), \\ E_1^\lambda &= \frac{1}{2} \rho^\lambda |u^\lambda|^2 + \rho^\lambda \varepsilon_1^\lambda, \end{aligned} \quad (2.1)$$

where $p_1(\rho^\lambda, \varepsilon_1^\lambda) = (\gamma_1 - 1)\rho^\lambda \varepsilon_1^\lambda$ with γ_1 a given constant greater than 1. One can prove [2] that the relaxation system (2.1) can be supplemented by entropy inequalities under the form

$$\partial_t \rho^\lambda \Sigma + \operatorname{div}(\rho^\lambda \Sigma u^\lambda) \leq \operatorname{RED}^\lambda := -\lambda \rho^\lambda (\Sigma_{,s_1 s_1 \varepsilon_1} - \Sigma_{,\varepsilon_2}) (\varepsilon_2 - \phi(\rho^\lambda, \varepsilon_1^\lambda)),$$

where $s_1(\rho, \varepsilon_1) = \rho^{\gamma_1 - 1} / \varepsilon_1$ and the specific entropy Σ denotes an arbitrary function in $C^1(\mathbb{R}_+^2)$ such that $\rho \Sigma$ is convex in $(\rho, \rho \varepsilon_1, \rho \varepsilon_2)$ and that it can be written under the form $\Sigma = \Sigma(s_1(\rho, \varepsilon_1), \varepsilon_2)$. $\operatorname{RED}^\lambda$ represents the rate of entropy dissipation.

Formally, the original Euler system (1.1) will be recovered at $\lambda \rightarrow +\infty$ with

$$\varepsilon = \varepsilon_1 + \varepsilon_2 = \varepsilon_1 + \phi(\rho, \varepsilon_1), \quad (2.2)$$

provided that we have the condition (called the *consistency condition*)

$$p(\rho, \varepsilon_1 + \phi(\rho, \varepsilon_1)) = p_1(\rho, \varepsilon_1) = (\gamma_1 - 1)\rho \varepsilon_1. \quad (2.3)$$

This last condition can be fulfilled for any given choice of $\gamma_1 > 1$.

But, in addition to the conservative system (1.1), one also wants to recover at the limit the entropy inequality (1.2). The following result, due to Coquel and Perthame [2], gives this last condition under a characterization of the admissible γ_1 .

THEOREM 1. *Assuming that γ_1 satisfies*

$$\begin{aligned} \gamma_1 &> \sup_{\rho, \varepsilon} \Gamma(\rho, \varepsilon), \quad \text{where } \Gamma(\rho, \varepsilon) = 1 + p_{\varepsilon}/\rho, \\ \gamma_1 &> \sup_{\rho, \varepsilon} \gamma(\rho, \varepsilon), \quad \text{where } \gamma(\rho, \varepsilon) = \frac{\rho}{p} p_{\varepsilon} + p_{\varepsilon}/\rho, \end{aligned} \quad (2.4)$$

provided that γ_1 is finite, we then have

(i) *there exists a (unique) specific entropy $\Sigma(s_1, \varepsilon_2)$ such that at equilibrium ($\varepsilon = \varepsilon_1 + \phi(\rho, \varepsilon_1)$)*

$$s(\rho, \varepsilon) = \Sigma(s_1(\rho, \varepsilon_1), \phi(\rho, \varepsilon_1)),$$

(ii) *this entropy is uniformly compatible with the relaxation procedure, i.e.*

$$RED^\lambda \leq 0 \quad \text{for all } \lambda > 0.$$

2.2. WENO Schemes

We use the fifth-order WENO scheme in [10]. For a scalar conservation law

$$u_t + f(u)_x = 0; \quad (2.5)$$

the derivative $f(u)_x$ at the grid point $x = x_j$ is approximated by a conservative flux difference,

$$f(u)_x|_{x=x_j} \approx \frac{1}{\Delta x} (\hat{f}_{j+1/2} - \hat{f}_{j-1/2}). \quad (2.6)$$

The WENO numerical flux $\hat{f}_{j+1/2}$ is computed as follows. For a positive wind direction $f'(u) \geq 0$, we first define three third-order numerical fluxes:

$$\begin{aligned} \hat{f}_{j+1/2}^1 &= \frac{1}{3}f(u_{j-2}) - \frac{7}{6}f(u_{j-1}) + \frac{11}{6}f(u_j), \\ \hat{f}_{j+1/2}^2 &= -\frac{1}{6}f(u_{j-1}) + \frac{5}{6}f(u_j) + \frac{1}{3}f(u_{j+1}), \\ \hat{f}_{j+1/2}^3 &= \frac{1}{3}f(u_j) + \frac{5}{6}f(u_{j+1}) - \frac{1}{6}f(u_{j+2}). \end{aligned} \quad (2.7)$$

A third-order ENO scheme will result if we choose one of the three third-order fluxes in (2.7) adequately, according to the size of divided differences [19]. On the other hand, a fifth-order linear scheme will result if we choose the flux as

$$\hat{f}_{j+1/2}^{\text{linear}} = d_1 \hat{f}_{j+1/2}^1 + d_2 \hat{f}_{j+1/2}^2 + d_3 \hat{f}_{j+1/2}^3 \quad (2.8)$$

with

$$d_1 = \frac{1}{10}, \quad d_2 = \frac{3}{5}, \quad d_3 = \frac{3}{10}. \quad (2.9)$$

The fifth-order WENO scheme results if we choose the numerical flux as

$$\hat{f}_{j+1/2} = \omega_1 \hat{f}_{j+1/2}^1 + \omega_2 \hat{f}_{j+1/2}^2 + \omega_3 \hat{f}_{j+1/2}^3, \quad (2.10)$$

with ω_i defined by

$$\omega_i = \frac{\alpha_i}{\sum_{s=1}^3 \alpha_s}, \quad \alpha_i = \frac{d_i}{(\epsilon + \beta_i)^2}, \quad (2.11)$$

and

$$\begin{aligned} \beta_1 &= \frac{13}{12}(f(u_{i-2}) - 2f(u_{i-1}) + f(u_i))^2 + \frac{1}{4}(f(u_{i-2}) - 4f(u_{i-1}) + 3f(u_i))^2, \\ \beta_2 &= \frac{13}{12}(f(u_{i-1}) - 2f(u_i) + f(u_{i+1}))^2 + \frac{1}{4}(f(u_{i-1}) - f(u_{i+1}))^2, \\ \beta_3 &= \frac{13}{12}(f(u_i) - 2f(u_{i+1}) + f(u_{i+2}))^2 + \frac{1}{4}(3f(u_i) - 4f(u_{i+1}) + f(u_{i+2}))^2. \end{aligned} \quad (2.12)$$

In all our numerical examples ϵ in (2.11) is taken as 10^{-6} , as was done in [10]. The weights in (2.11) are chosen so that in smooth regions (including at smooth extrema), the WENO flux (2.10) behaves similarly to the linear flux (2.8) and is uniformly fifth-order accurate. Near shocks, however, the WENO flux (2.10) behaves similarly to an ENO flux, in the sense that any stencil crossing a discontinuity has a near-zero weight. For details of the derivation see [10, 21].

If the wind direction is negative, $f'(u) \leq 0$, the procedure is symmetric to the case with $f'(u) \geq 0$, with respect to the location $x_{j+1/2}$. In general, a flux splitting is used,

$$f(u) = f^+(u) + f^-(u), \quad (2.13)$$

such that $f^+(u)$ has a positive wind direction and $f^-(u)$ has a negative wind direction:

$$\frac{d}{du} f^+(u) \geq 0, \quad \frac{d}{du} f^-(u) \leq 0. \quad (2.14)$$

The procedure described above can then be applied to $f^+(u)$ and $f^-(u)$ separately. The simplest flux splitting is the Lax–Friedrichs splitting,

$$f^\pm(u) = \frac{1}{2}(f(u) \pm \alpha u), \quad (2.15)$$

where

$$\alpha = \max_u |f'(u)|.$$

Notice that while first- and second-order schemes with a Lax–Friedrichs splitting are quite dissipative, higher order schemes based on the Lax–Friedrichs fluxes usually give very good results. We use Lax–Friedrichs fluxes in this paper.

For systems of conservation laws in this paper, we use both a component-wise version, where the procedure described above is applied to each equation in the system separately,

and a characteristic version, where locally we apply the procedure described above on characteristic projections. For details about how to perform a local characteristic procedure see, for example, [18, 19, 10].

2.3. Construction of the Relaxed WENO Scheme

The procedure to solve the Euler system (1.1) within the framework of the energy relaxation theory is the following. Given the numerical equilibrium solution at the time level t^n

$$\rho(x, t^n), u(x, t^n), \varepsilon(x, t^n), \quad (2.16)$$

this approximation is advanced to the next time level $t^{n+1} = t^n + \Delta t$ in two steps:

- First step: relaxation. The two internal energies $\varepsilon_1(x, t^n)$ and $\varepsilon_2(x, t^n)$ are obtained by (2.2) and the consistency condition (2.3):

$$\begin{aligned} \varepsilon_1(x, t^n) &= \frac{p(\rho(x, t^n), \varepsilon(x, t^n))}{(\gamma_1 - 1)\rho(x, t^n)}, \\ \varepsilon_2(x, t^n) &= \varepsilon(x, t^n) - \varepsilon_1(x, t^n). \end{aligned} \quad (2.17)$$

Notice that this step involves just one call to the pressure law per grid point and does not involve any derivatives of the pressure law or any iterations.

- Second step: evolution in time. For $t^n \leq t \leq t^{n+1}$, we solve the Cauchy problem for the relaxation system (2.1), with zero on the right side,

$$\begin{aligned} \partial_t \rho^\lambda + \operatorname{div}(\rho^\lambda u^\lambda) &= 0, \quad t \geq 0, x \in \mathbb{R}^d, \\ \partial_t \rho^\lambda u^\lambda + \operatorname{div}(\rho^\lambda u^\lambda \otimes u^\lambda + p_1^\lambda) &= 0, \\ \partial_t E_1^\lambda + \operatorname{div}((E_1^\lambda + p_1^\lambda)u^\lambda) &= 0, \\ \partial_t \rho^\lambda \varepsilon_2^\lambda + \operatorname{div}(\rho^\lambda u^\lambda \varepsilon_2^\lambda) &= 0, \\ E_1^\lambda &= \frac{1}{2} \rho^\lambda |u^\lambda|^2 + \rho^\lambda \varepsilon_1^\lambda, \end{aligned} \quad (2.18)$$

and the initial data,

$$\rho(x, t^n), u(x, t^n), \varepsilon_1(x, t^n), \varepsilon_2(x, t^n), \quad (2.19)$$

and we obtain at time t^{n+1-} ,

$$\rho(x, t^{n+1-}), u(x, t^{n+1-}), \varepsilon_1(x, t^{n+1-}), \varepsilon_2(x, t^{n+1-}). \quad (2.20)$$

At last, we compute the equilibrium solution at time t^{n+1} by

$$\begin{aligned} \rho(x, t^{n+1}) &= \rho(x, t^{n+1-}), \\ u(x, t^{n+1}) &= u(x, t^{n+1-}), \\ \varepsilon(x, t^{n+1}) &= \varepsilon_1(x, t^{n+1-}) + \varepsilon_2(x, t^{n+1-}). \end{aligned} \quad (2.21)$$

Remark 1. The first step is clearly a relaxation phase, as it is equivalent to the solution of the ODE problem for $t \geq t^n$,

$$\begin{aligned} d_t \rho^\lambda &= 0, \\ d_t \rho^\lambda u^\lambda &= 0, \\ d_t E_1^\lambda &= \lambda \rho^\lambda (\varepsilon_2^\lambda - \phi(\rho^\lambda, \varepsilon_1^\lambda)), \\ d_t \rho^\lambda \varepsilon_2^\lambda &= -\lambda \rho^\lambda (\varepsilon_2^\lambda - \phi(\rho^\lambda, \varepsilon_1^\lambda)), \end{aligned} \quad (2.22)$$

with initial data at time level t^n ,

$$\rho(x, t^{n-}), u(x, t^{n-}), \varepsilon_1(x, t^{n-}), \varepsilon_2(x, t^{n-}). \quad (2.23)$$

and we let $\lambda \rightarrow +\infty$.

We now describe the numerical method we will use for the step of evolution in time. Although our numerical results concern both one- and two-dimensional problems, for simplicity of presentation we shall restrict our description to one space dimension. As we are using the finite difference version of WENO schemes in [10], extensions to two and more spatial dimensions are simply done dimension by dimension. Essentially, the two-dimensional code is the one-dimensional code with an outside “do loop.”

We have to solve for $t^n \leq t < t^{n+1}$ the system of four equations,

$$\partial_t U + \partial_x F(U) = 0, \quad + \text{initial conditions given by (2.19)}, \quad (2.24)$$

where

$$\begin{aligned} U &= (\rho, \rho u, E_1, \rho \varepsilon_2)^T, \\ F(U) &= (\rho u, \rho u^2 + p_1, (E_1 + p_1)u, \rho u \varepsilon_2)^T. \end{aligned} \quad (2.25)$$

In order to solve the ordinary differential equation

$$\frac{d}{dt} U = L(U), \quad (2.26)$$

where $L(U)$ is a discretization of the spatial operator, we use a third-order TVD Runge–Kutta scheme [18].

Remark 2. We have two possibilities for the placement of the relaxation step: each Runge–Kutta inner stage or each time step. With the first example of section 3.3, we show that the two approaches give nearly identical results in accuracy. Of course the second approach is less costly. We thus perform all our calculations using the second approach.

We now discretize the space into uniform intervals of size Δx and denote $x_j = j \Delta x$. Various quantities at x_j will be identified by the subscript j .

We use the WENO procedure described in the previous subsection to obtain the spatial operator $L_j(U)$ which approximates $-\partial_x F(U)$ at x_j . We have tested several possibilities for the definition of $L(U)$ based on WENO schemes. The first one is to use a WENO Lax–Friedrichs scheme with a full characteristic decomposition. For this purpose we need

to compute a Roe matrix for the system (2.24) and its eigenvalues and eigenvectors. The details of this derivation are included in Appendix A.

The other possibility is to compute the first three components of the numerical flux $\hat{F}_{j+1/2}^1$, $\hat{F}_{j+1/2}^2$, $\hat{F}_{j+1/2}^3$ by using a WENO Lax–Friedrichs scheme with a decomposition on the Euler system characteristics and to obtain the last numerical flux $\hat{F}_{j+1/2}^4$ with a scalar WENO Lax–Friedrichs scheme. This is possible because the first three equations of system (2.24) are independent from the last one.

Remark 3. We have also tried to compute the last numerical flux by using a first-order scheme specially designed in order to preserve the maximum principle for ε_2 [11]. But with this approach, we lose the accuracy of the high-order WENO scheme also for the other variables.

Remark 4. In order to make comparisons in the numerical results we have also implemented a WENO Lax–Friedrichs scheme with a full characteristic decomposition for a two molecular vibrating gas (see next section for a description of the related EOS). For this purpose we need a definition of the corresponding Roe average matrix. We give it in Appendix B. For the numerical comparisons for the other real gases we use a component-wise WENO Lax–Friedrichs scheme which requires only the computation of the sound velocity

$$c = \sqrt{p_{,\rho} + p(p_{,\varepsilon}/\rho^2)}. \quad (2.27)$$

3. NUMERICAL RESULTS

3.1. Description of the Different Equations of States

We present here several equations of states which we will use in the computation. We find the second one in the paper of In [9], while the third one comes from Glaister [4]:

- *Polytropic ideal gas.* The equation of states for a polytropic ideal gas (also called perfect gas) is

$$p(\rho, \varepsilon) = (\gamma - 1)\rho\varepsilon. \quad (3.1)$$

Then we have

$$p_{,\rho} = (\gamma - 1)\varepsilon, \quad p_{,\varepsilon} = (\gamma - 1)\rho. \quad (3.2)$$

Air under normal conditions (p and T moderate enough) can be considered as a perfect gas with $\gamma = 7/5 = 1.4$ (approximately a mixture of two diatomic molecular species: 20% O_2 ; 80% N_2).

- *Two-molecular vibrating gas.* When the temperature increases the vibrational motion of oxygen and nitrogen molecules in air becomes important, and specific heats vary with temperatures. So that one must consider the thermally perfect, calorically imperfect model for two-molecular vibrating gas,

$$p(\rho, \varepsilon) = r\rho T(\varepsilon), \quad (3.3)$$

where the temperature T is given by the implicit expression

$$\rho\varepsilon = c_v^{\text{tr}} T + \rho \left(\frac{\alpha \Theta_{\text{vib}}}{\exp(\Theta_{\text{vib}}/T) - 1} \right), \quad (3.4)$$

with $r = 287.086 J \cdot kg^{-1} \cdot K^{-1}$, $C_v^{tr} = r/(\gamma_{tr} - 1)$, $\gamma_{tr} = 1.4$, $\Theta_{vib} = 10^3 K$, and $\alpha = r$. Then we have

$$p_{,\rho} = rT(\varepsilon), \quad p_{,\varepsilon} = \frac{r\rho}{\varepsilon'(T(\varepsilon))}. \quad (3.5)$$

• *Osborne model.* R. K. Osborne from the Los Alamos Scientific Laboratory has developed a quite general equation of states in the form [17]

$$p(\rho, \varepsilon) = \frac{1}{E + \phi_0} (\zeta(a_1 + a_2\zeta) + E(b_0 + \zeta(b_1 + b_2\zeta) + E(c_0 + c_1\zeta))), \quad (3.6)$$

where $E = \rho_0\varepsilon$ and $\zeta = \rho/\rho_0 - 1$ and the constants ρ_0 , a_1 , a_2 , b_0 , b_1 , b_2 , c_0 , c_1 , ϕ_0 depend on the material in question. The typical values for water are $\rho_0 = 10^{-2}$, $a_1 = 3.84 \times 10^{-4}$, $a_2 = 1.756 \times 10^{-3}$, $b_0 = 1.312 \times 10^{-2}$, $b_1 = 6.265 \times 10^{-2}$, $b_2 = 0.2133$, $c_0 = 0.5132$, $c_1 = 0.6761$, and $\phi_0 = 2. \times 10^{-2}$. Then we have

$$\begin{aligned} p_{,\rho} &= \frac{1}{\rho_0(E + \phi_0)} ((a_1 + 2a_2\zeta) + E(b_1 + 2b_2\zeta + Ec_1)), \\ p_{,\varepsilon} &= -\frac{\rho_0}{E + \phi_0} p + \frac{\rho_0}{E + \phi_0} (b_0 + \zeta(b_1 + b_2\zeta) + 2E(c_0 + c_1\zeta)). \end{aligned} \quad (3.7)$$

3.2. One-Dimensional Cases

EXAMPLE 1 (1D Riemann problems with perfect gas). We consider here two well-known problems which have the following Riemann type initial conditions:

$$\mathbf{u}(x, 0) = \begin{cases} \mathbf{u}_L & \text{if } x < 0, \\ \mathbf{u}_R & \text{if } x > 0. \end{cases}$$

The first one is Sod's problem [22]. The initial data are

$$(\rho_L, u_L, p_L) = (1, 0, 1), \quad (\rho_R, u_R, p_R) = (0.125, 0, 0.1).$$

The second one is proposed by Lax [12] with

$$(\rho_L, u_L, p_L) = (0.445, 0.698, 3.528), \quad (\rho_R, u_R, p_R) = (0.5, 0, 0.571).$$

Of course, for this perfect gas situation there is no need to use the relaxation model in practice. The purpose of this test problem is to test the behavior of different relaxation models (different γ_1 's) and different ways of treating the relaxed system (fully characteristic and partially characteristic for the first three equations only).

For this example, a uniform grid of 100 points are used and every 2 points are drawn in the figures.

We first give, in Table I, a CPU time comparison among the traditional WENO characteristic scheme for the perfect gas, and the WENO scheme applied to the relaxation system, both with a fully characteristic decomposition and with a partially characteristic decomposition for the first three equations only. The calculation is done on a SUN Ultra1 workstation. We can see that, while a fully characteristic decomposition is significantly more costly, the

TABLE I
CPU Time (in seconds) of Different Schemes for the Sod and Lax Shock Tube Problems for a Perfect Gas

Case	WENO with characteristic	Relaxed WENO with full characteristic	Relaxed WENO with partial characteristic
Sod shock	2.28	3.49	2.91
Lax shock	3.32	4.93	4.08

partially characteristic decomposition is only slightly more costly than the WENO scheme applied to the original perfect gas Euler equations.

In Figs. 1 and 3, we present the comparison for the Sod's and Lax's shock tube problems, of the fifth-order WENO schemes, applied directly to the perfect gas Euler equations using a characteristic decomposition and applied to the relaxation model with $\gamma_1 = 3$, using only partial characteristic decomposition of the first three equations. We can see that the results are very close, except for the slight over- and undershoots in entropy for the relaxation model calculation. This indicates the feasibility of using the relaxation model.

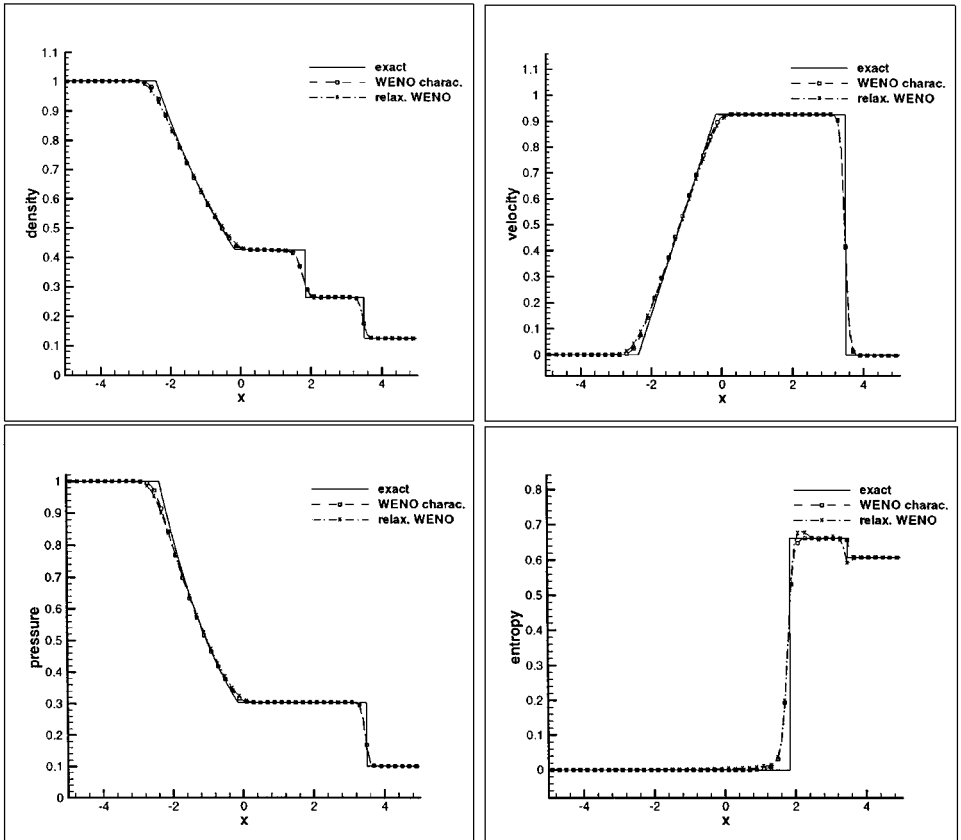


FIG. 1. Sod's shock tube problem with WENO-LF-5 characteristic and relaxed WENO-LF-5 partial characteristic with $\gamma_1 = 3.0$.

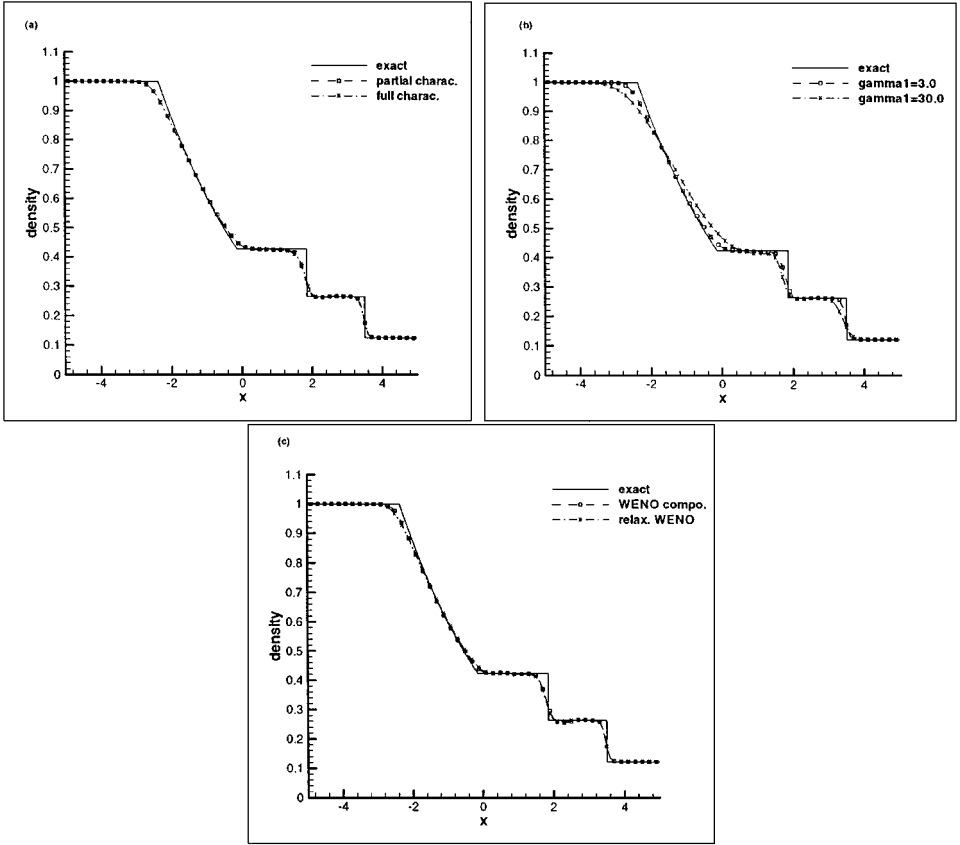


FIG. 2. Sod's shock tube problem with WENO-LF-5. Comparisons of partial and full characteristic decompositions for the relaxation model with $\gamma_1 = 3$ (a), $\gamma_1 = 3$ and $\gamma_1 = 30$ for the relaxation model with partial characteristic decomposition (b), and the relaxation model with partial characteristic decomposition with $\gamma_1 = 3$ versus the component-wise WENO applied to the original perfect gas Euler equations (c).

In Figs. 2 and 4, we present the comparison for the Sod's and Lax's shock tube problems, of the fifth-order WENO schemes. The top left figure compares the full characteristic decomposition for the relaxation model, with a partial characteristic decomposition for the first three equations only, for $\gamma_1 = 3$. We can see that the results are quite close, again indicating the feasibility of using the less costly partial characteristic decomposition for the relaxation model. The top right figure compares the effect of different γ_1 's in the relaxation model. The apparently bigger γ_1 corresponds to larger numerical dissipation. This indicates that one should always choose the smallest possible γ_1 , subject to stability considerations. The bottom figure compares the relaxation WENO results for $\gamma_1 = 3$ and a partial characteristic decomposition, with a component-wise WENO scheme applied directly on the original perfect gas Euler equations. Although neither uses the correct characteristic information, apparently the relaxation model results are better than the componentwise results, especially for the Lax's problem in Fig. 4.

EXAMPLE 2 (1D Riemann problems with real gases). In this example we compute the solutions to the Riemann shock tube problem for the two-molecular vibrating gas (3.3)–(3.5) and the Osborne model (3.6)–(3.7), with the following initial conditions in Table II.

TABLE II
Initial Conditions for the Test Cases for Real Gases

Case	State	Density	Velocity	Specific internal energy
A	Left	0.066	0.0	7.22e6
	Right	0.030	0.0	1.44e6
B	Left	1.40	0.0	2.22e6
	Right	0.14	0.0	2.24e6
C	Left	1.2900	0.0	1.95e6
	Right	0.0129	0.0	2.75e6
D	Left	1.00	0.0	2.00e6
	Right	0.01	0.0	2.50e5
E	Left	0.01	2200.0	1.44e5
	Right	0.14	0.0	4.00e5

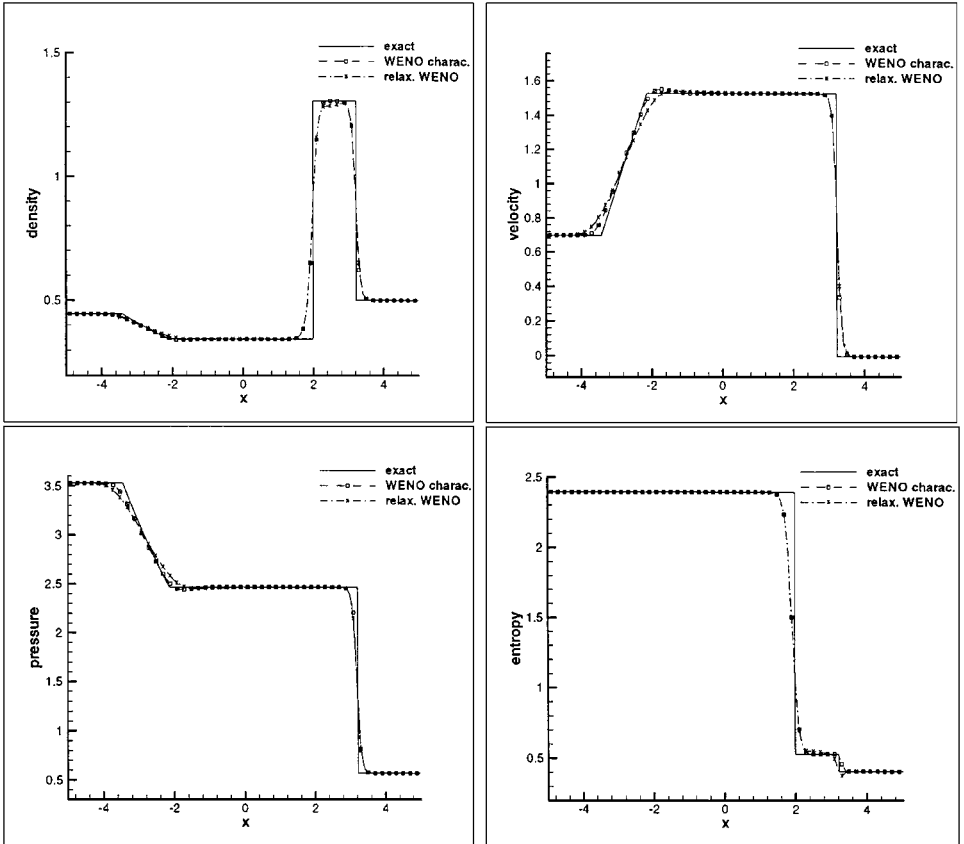


FIG. 3. Lax's shock tube problem with WENO-LF-5 characteristic and relaxed WENO-LF-5 partial characteristic with $\gamma_1 = 3.0$.

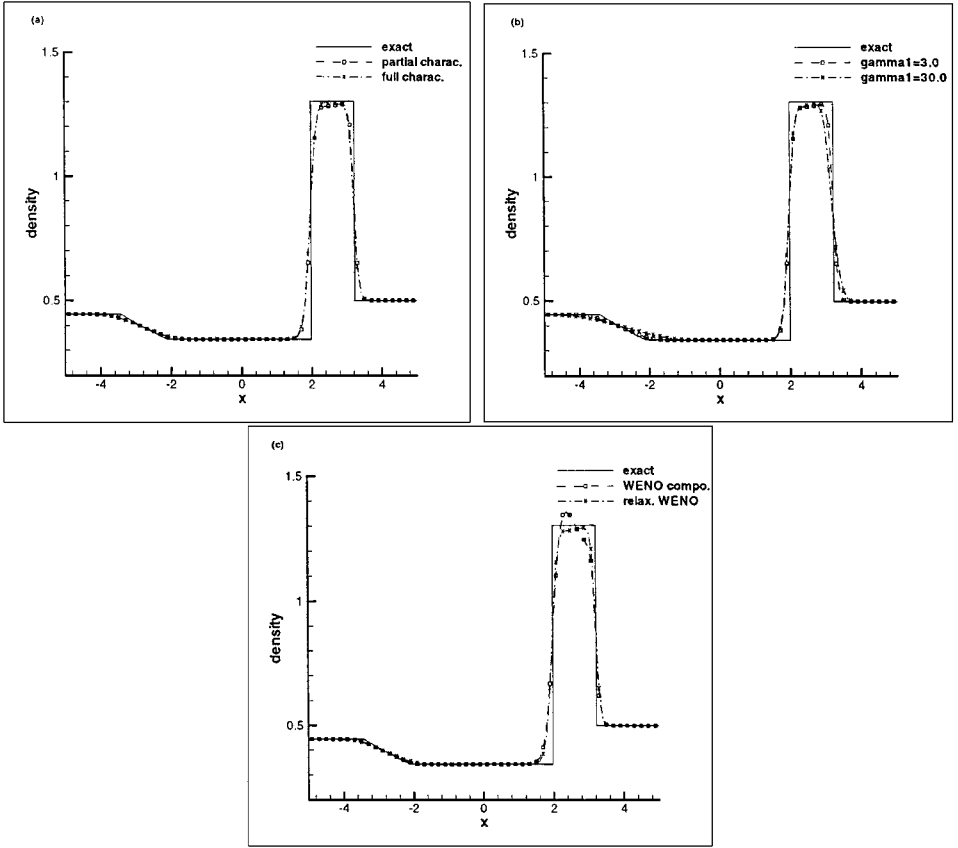


FIG. 4. Lax’s shock tube problem with WENO-LF-5. Comparisons of partial and full characteristic decompositions for the relaxation model with $\gamma_1 = 3$ (a), $\gamma_1 = 3$ and $\gamma_1 = 30$ for the relaxation model with partial characteristic decomposition (b), and the relaxation model with partial characteristic decomposition with $\gamma_1 = 3$ versus the component-wise WENO applied to the original perfect gas Euler equations (c).

For this example, a uniform grid of 200 points are used and every four points are drawn in the figures. Also, the “exact solutions” in the figures are obtained with the best scheme, using 2000 points.

We first give a CPU time comparison between the full characteristic decomposition for the original model and the partial characteristic decomposition, using only the first three equations of the relaxation model, for the two-molecular vibrating gas model, in Table III. We can see that the partial characteristic decomposition for the relaxed model is usually less than half as costly than the full characteristic version for the original system. Although the relaxed model has one more equation, it does not require the computation of the complicated derivatives of the EOS.

In Fig. 5 we show the comparison of the full characteristic decomposition for the original model and the partial characteristic decomposition using only the first three equations of the relaxation model, for the two-molecular vibrating gas model, with the case A initial condition. The results are almost identical, indicating that the relaxation model with a partial characteristic decomposition works well with a much-reduced cost.

In Fig. 6 we show the comparison of the component WENO scheme on the original system and the partially characteristic WENO scheme on the relaxed system with $\gamma_1 = 2.0$,

TABLE III
CPU Time (in Seconds) Depending on Full or Partial Characteristic Decomposition with a Two-Molecular Vibrating Gas

Case	WENO with characteristic	Relaxed WENO with partial characteristic
A	12.68	5.21
B	4.8	2.63
C	12.53	4.87
D	15.0	5.35
E	15.0	7.84

for the Osborne gas model with the case A initial condition. We can see that the result of the relaxed model is much better, especially for the density. This indicates that the relaxation model is a good one for the computation of real gases.

In Fig. 7 we show the comparison of taking $\gamma_1 = 10$, which satisfies the stability condition (2.4), and $\gamma_1 = 2$, which satisfies only the second inequality in the stability condition (2.4) for the partial characteristic decomposition using only the first three equations of the relaxation

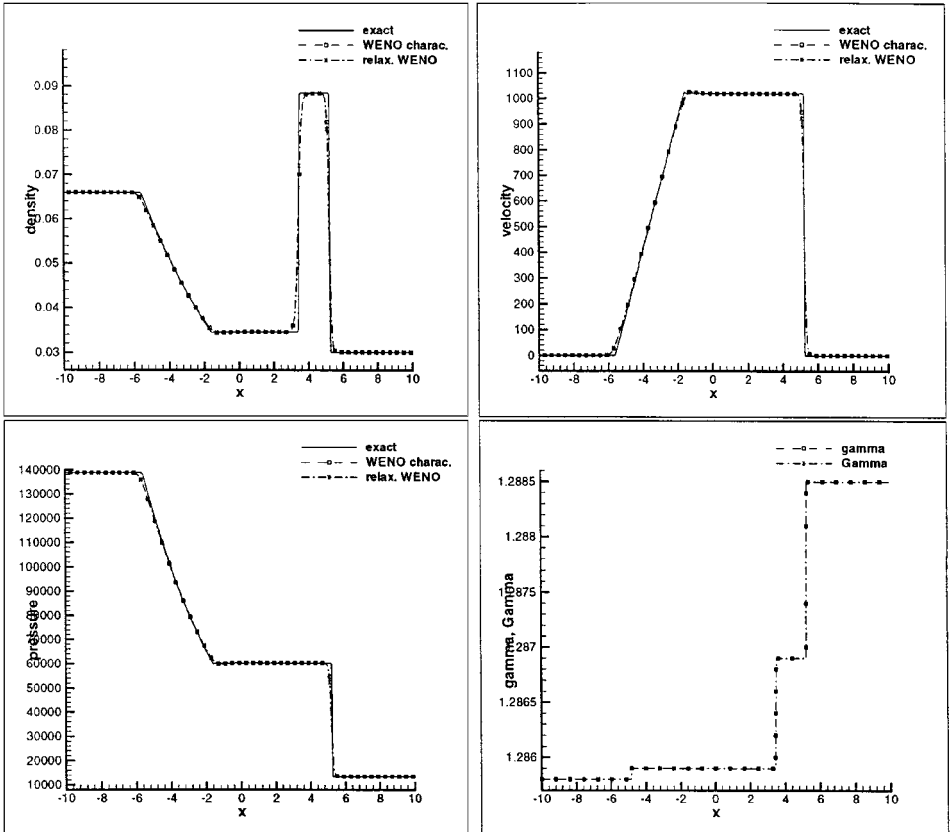


FIG. 5. Case A + two vibrating molecular gas model with WENO-LF-5 characteristic and relaxed WENO-LF-5 partial characteristic with $\gamma_1 = 1.5$.

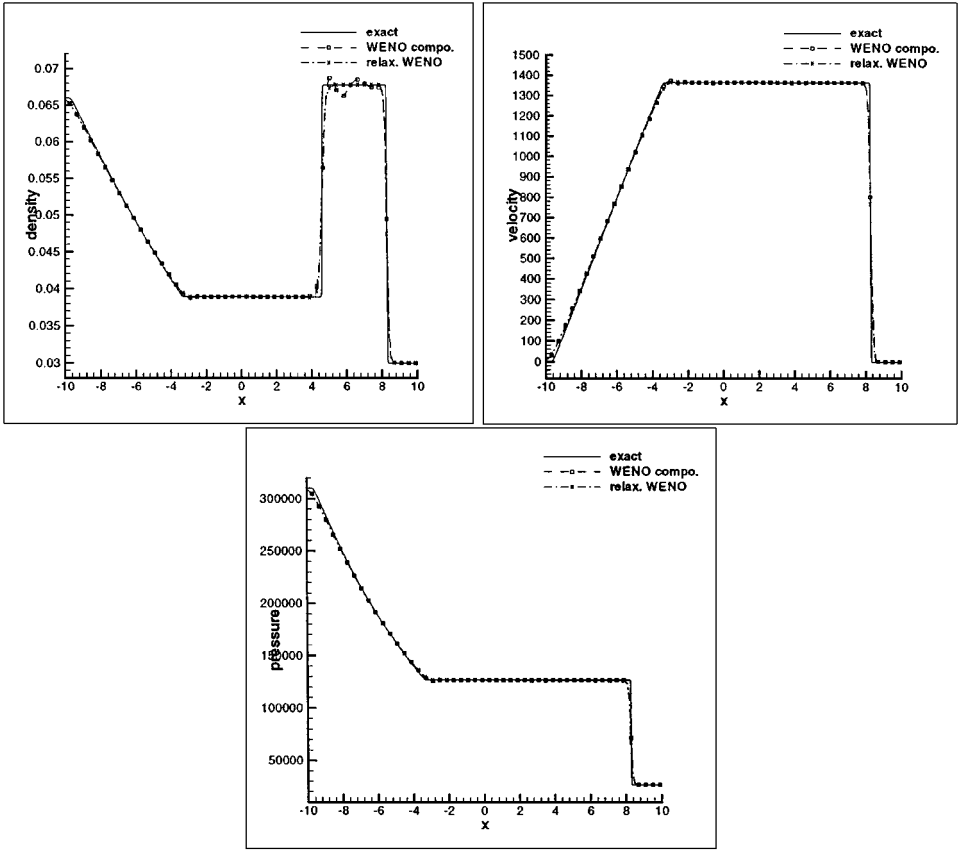


FIG. 6. Case A + Osborne gas model with componentwise WENO-LF-5 for the original system and relaxed WENO-LF-5 partial characteristic with $\gamma_1 = 2.0$.

model and the Osborne gas model with the case A initial condition. We can see that the $\gamma_1 = 2$ results are stable and less dissipative, indicating that in practice one does not always have to choose γ_1 satisfying both inequalities in condition (2.4).

We have also tested the same problems for the other initial condition cases B, C, D, and E. The results are mostly similar qualitatively, as in case A. To save space we will not present the results here.

3.3. Two-Dimensional Cases

EXAMPLE 3 (An isentropic vortex). This example is used to verify the accuracy of the relaxation approach, especially the placement of the relaxation steps during time stepping. The gas is ideal but we still use the relaxation model. We consider the following idealized problem for the Euler equations in 2D: the mean flow is $\rho = 1$, $p = 1$ and $(u, v) = (1, 1)$ (diagonal flow). We add to this flow an isentropic vortex (perturbation in (u, v) and the temperature $T = p/\rho$; no perturbation in the entropy $S = P/\rho^\gamma$),

$$(\delta u, \delta v) = \frac{\varepsilon}{2\pi} \exp\left(\frac{1-r^2}{2}\right) (-\bar{y}, \bar{x}), \quad \delta T = -\frac{(\gamma-1)\varepsilon^2}{8\gamma\pi^2} \exp(1-r^2), \quad \delta S = 0,$$

where $(\bar{x}, \bar{y}) = (x-5, y-5)$, $r^2 = \bar{x}^2 + \bar{y}^2$, and the vortex strength $\varepsilon = 5$. See [21].

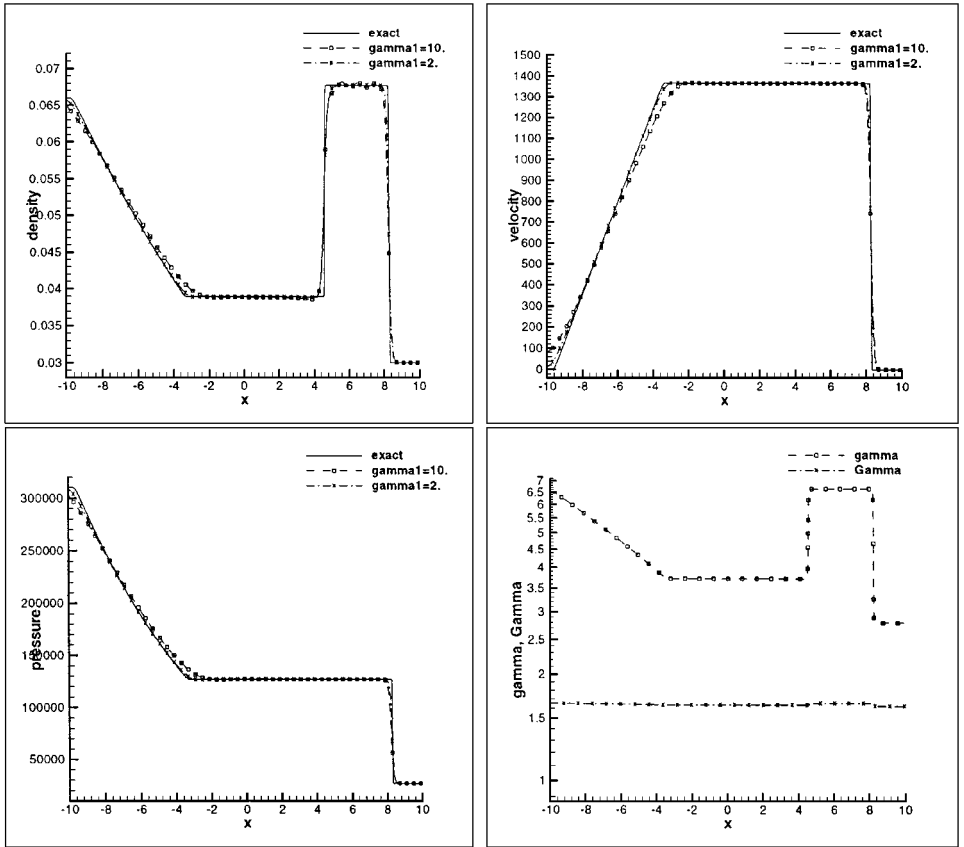


FIG. 7. Case A + Osborne gas model with the relaxed WENO-LF-5 partial characteristic with $\gamma_1 = 10.0$ and $\gamma_1 = 2.0$.

The computational domain is taken as $[0, 10,] \times [0, 10]$, extended periodically in both directions. This allows us to perform long-time simulation without having to deal with a large domain.

It is clear that the exact solution of the Euler equation with the above initial and boundary conditions is just the passive convection of the vortex with the mean velocity.

In Table IV we show the accuracy result at $t = 10$ (one time period). We can see that WENO for the relaxed model with $\gamma_1 = 3$ gives a somewhat larger error than WENO applied directly to the original system, but the order of accuracy is correct. Moreover, to place the relaxation step for each Runge–Kutta inner stage or just for each time step seems to give

TABLE IV
L1 Error and Order of Accuracy at $t = 10$ (1 Period)

Nb. points	WENO		Relaxed WENO each time step		Relaxed WENO each R-K step	
	L1 error	Accuracy	L1 error	Accuracy	L1 error	Accuracy
20×20	1.07e-2		1.22e-2		1.22e-2	
40×40	1.06e-3	3.3	2.16e-3	2.5	2.17e-3	2.5
80×80	6.50e-5	4.0	1.77e-4	3.6	1.78e-4	3.6
160×160	2.09e-6	4.9	7.57e-6	4.6	7.60e-6	4.6

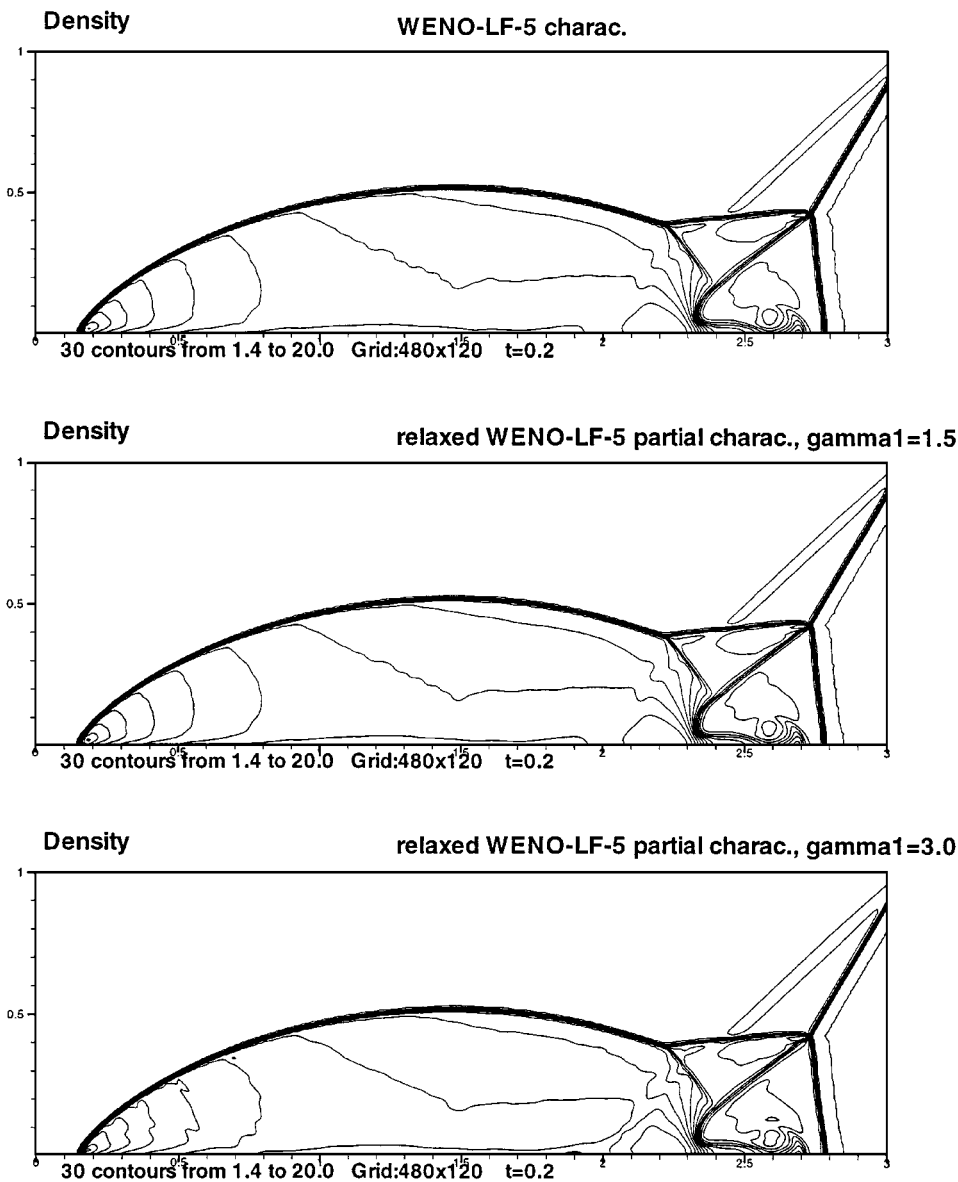


FIG. 8. Double-Mach reflection, perfect gas, 480×120 grid points.

almost identical results. We have thus used the less costly version of putting the relaxation step for every time step in all the numerical examples in this paper.

EXAMPLE 4 (Double Mach reflection). The computational domain is chosen to be $[0, 4] \times [0, 1]$, although only part of it ($[0, 3] \times [0, 1]$) is shown. The reflecting wall lies at the bottom of the computational domain, starting from $x = 1/6$. Initially a right-moving Mach 10 shock is positioned at $(x, y) = (1/6, 0)$ and makes a 60° angle with the x axis. For the bottom boundary, the exact postshock condition is imposed for the part from $x = 0$ to $x = 1/6$ and a reflective boundary condition is used for the rest. At the top boundary of our computational domain, the flow values are set to describe the exact motion of the Mach 10 shock (See [23] for a detailed description of this problem.)

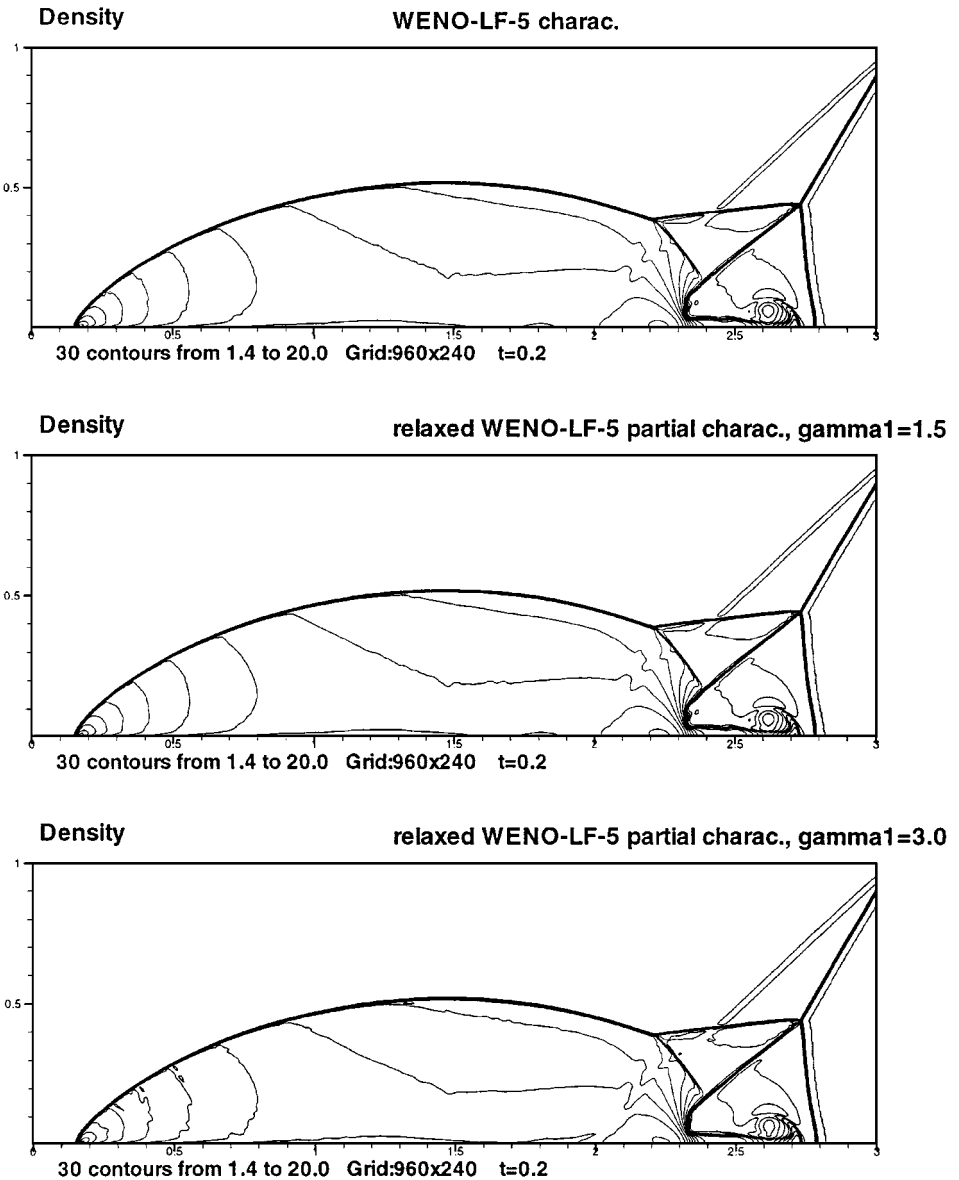


FIG. 9. Double-Mach reflection, perfect gas, 960×240 grid points.

First we present the results for a perfect gas. We compare the results using WENO directly on the original system [10] and using it on the relaxed model with $\gamma_1 = 1.5$ and $\gamma_1 = 3.0$ in Fig. 8 for a mesh of 480×120 points and Fig. 9 for a mesh of 960×240 points. We can see that the relaxed model results are quite satisfactory, although a bigger γ_1 results in some small oscillations.

Next, we show the results of the same problem with the two-molecular vibrating gas. The purpose here is to show that the relaxation model-based algorithm does work, rather than the details of the flow with more physical models. The results with both a 480×120 grid and a 960×240 grid are shown in Fig. 10. Comparing with the results in [3], we can

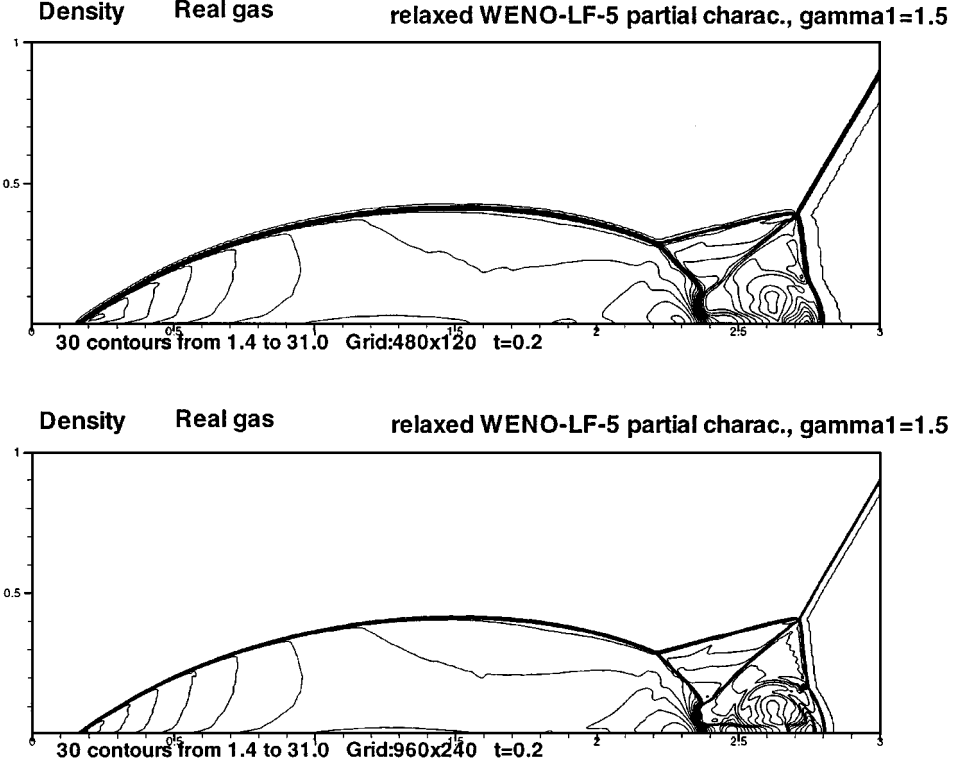


FIG. 10. Double-Mach reflection, two vibrating molecular gas.

see that the main features, such as the main shock being closer to the bottom boundary and the shock below the triple point being bent, are also observed here.

4. CONCLUDING REMARKS

We have applied the fifth-order WENO schemes to a relaxation model to compute the Euler equations of real gases. The algorithm does not depend on the specific form of the equation of states and does not need to compute the derivatives of the pressure law. One- and two-dimensional examples are shown to illustrate the accuracy and robustness of the algorithm. The algorithm seems to work well, even for strong (Mach 10) shocks. The choice of γ_1 in the relaxation model influences the numerical results, as the bigger γ_1 implies larger numerical viscosity. In practice one should choose γ_1 as small as possible, subject to stability restraints.

APPENDIX A: ROE MATRIX FOR THE RELAXATION SYSTEM

Let us consider two states U_l and U_r ; then the Roe matrix for the relaxation system (2.24) is

$$\bar{A}(U_l, U_r) = \begin{pmatrix} 0 & 1 & 0 & 0 \\ (\gamma_1 - 3)\frac{\bar{u}^2}{2} & -(\gamma_1 - 3)\bar{\rho}\bar{u} & (\gamma_1 - 1) & 0 \\ \bar{u}(-\bar{H}_1 + (\gamma_1 - 1)\frac{\bar{u}^2}{2}) & \bar{H}_1 - (\gamma_1 - 1)\bar{u}^2 & \gamma_1\bar{u} & 0 \\ -\bar{\varepsilon}_2\bar{u} & \bar{\varepsilon}_2 & 0 & \bar{u} \end{pmatrix}, \quad (\text{A.1})$$

where the averaged state $\bar{\rho}$, \bar{u} , \bar{H}_1 are defined by

$$\begin{aligned}\bar{\rho} &= \sqrt{\rho_l} \sqrt{\rho_r}, & \bar{u} &= \alpha_l u_l + \alpha_r u_r, \\ \bar{H}_1 &= \alpha_l H_{1l} + \alpha_r H_{1r}, & \bar{\varepsilon}_2 &= \alpha_l \varepsilon_{1l} + \alpha_r \varepsilon_{1r},\end{aligned}\quad (\text{A.2})$$

with

$$\begin{aligned}H_1 &= (e_1 + p_1)/\rho, \\ \alpha_l &= \frac{\sqrt{\rho_l}}{\sqrt{\rho_l} + \sqrt{\rho_r}}, & \alpha_r &= 1 - \alpha_l = \frac{\sqrt{\rho_r}}{\sqrt{\rho_l} + \sqrt{\rho_r}}.\end{aligned}\quad (\text{A.3})$$

The four eigenvalues of \bar{A} are

$$\bar{a}_1 = \bar{u} - \bar{c}, \quad \bar{a}_2 = \bar{u}, \quad \bar{a}_3 = \bar{u} + \bar{c}, \quad \bar{a}_4 = \bar{u}, \quad (\text{A.4})$$

where the averaged sound speed \bar{c} has the usual expression

$$\bar{c} = \sqrt{\gamma_1 \bar{p}_1 / \bar{\rho}}. \quad (\text{A.5})$$

A set of right eigenvectors can be

$$\bar{r}_1 = \begin{pmatrix} 1 \\ \bar{u} - \bar{c} \\ \bar{H} - \bar{u}\bar{c} \\ \bar{\varepsilon}_2 \end{pmatrix}, \quad \bar{r}_2 = \begin{pmatrix} 1 \\ \bar{u} \\ \frac{\bar{u}^2}{2} \\ 0 \end{pmatrix}, \quad \bar{r}_3 = \begin{pmatrix} 1 \\ \bar{u} + \bar{c} \\ \bar{H} + \bar{u}\bar{c} \\ \bar{\varepsilon}_2 \end{pmatrix}, \quad \bar{r}_4 = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix}. \quad (\text{A.6})$$

And the corresponding orthogonal set of left eigenvectors is

$$\bar{l}_1 = \begin{pmatrix} \frac{1}{2}(b_1 + \frac{\bar{u}}{\bar{c}}) \\ -\frac{1}{2}(\bar{u}b_2 + \frac{1}{\bar{c}}) \\ \frac{b_2}{2} \\ 0 \end{pmatrix}, \quad \bar{l}_2 = \begin{pmatrix} 1 - b_1 \\ b_2 \bar{u} \\ -b_2 \\ 0 \end{pmatrix}, \quad \bar{l}_3 = \begin{pmatrix} \frac{1}{2}(b_1 - \frac{\bar{u}}{\bar{c}}) \\ -\frac{1}{2}(\bar{u}b_2 - \frac{1}{\bar{c}}) \\ \frac{b_2}{2} \\ 0 \end{pmatrix}, \quad \bar{l}_4 = \begin{pmatrix} -\bar{\varepsilon}_2 b_1 \\ \bar{\varepsilon}_2 \bar{u} b_2 \\ -\bar{\varepsilon}_2 b_2 \\ 1 \end{pmatrix}, \quad (\text{A.7})$$

where

$$b_1 = \frac{(\gamma_1 - 1)\bar{u}^2}{2\bar{c}^2}, \quad b_2 = \frac{(\gamma_1 - 1)}{\bar{c}^2}. \quad (\text{A.8})$$

APPENDIX B: ROE MATRIX FOR A TWO-MOLECULAR VIBRATING GAS

Let us consider two states U_l and U_r ; then the Roe matrix for an Euler system of real gas is (see [6] for details)

$$\bar{A}(U_l, U_r) = \begin{pmatrix} 0 & 1 & 0 \\ \bar{\chi} + \bar{\kappa}\bar{u}^2/2 & (2 - \bar{\kappa})\bar{u} & \bar{\kappa} \\ \bar{u}(\bar{\chi} + \bar{\kappa}\bar{u}^2/2 - \bar{H}) & \bar{H} - \bar{\kappa}\bar{u}^2 & (1 + \bar{\kappa})\bar{u} \end{pmatrix}, \quad (\text{B.1})$$

where \bar{u} , \bar{H} are the Roe average values of the velocity and the total specific enthalpy ($H = \varepsilon + 1/2u^2 + p/\rho$),

$$\bar{u} = \frac{\sqrt{\rho_l}u_l + \sqrt{\rho_r}u_r}{\sqrt{\rho_l} + \sqrt{\rho_r}}, \quad (\text{B.2})$$

$$\bar{H} = \frac{\sqrt{\rho_l}H_l + \sqrt{\rho_r}H_r}{\sqrt{\rho_l} + \sqrt{\rho_r}}, \quad (\text{B.3})$$

and $\bar{\chi}$ and $\bar{\kappa}$ are two parameters which must satisfy

$$\Delta p = \bar{\kappa} \Delta \rho \varepsilon + \bar{\chi} \Delta \rho \quad (\text{B.4})$$

with $\Delta \rho = \rho_r - \rho_l$, $\Delta \rho \varepsilon = \rho_r \varepsilon_r - \rho_l \varepsilon_l$, and $\Delta p = p(\rho_r, \varepsilon_r) - p(\rho_l, \varepsilon_l)$.

The definitions for $\bar{\kappa}$ and $\bar{\chi}$ proposed by In [9] for a two-molecular vibrating gas are

$$\bar{\kappa} = \begin{cases} \frac{r(T(\varepsilon_r) - T(\varepsilon_l))}{\varepsilon_r - \varepsilon_l} & \text{if } \varepsilon_r \neq \varepsilon_l, \\ \frac{1}{2} \left(\frac{p_{,\varepsilon}(\rho_l, \varepsilon)}{\rho_l} + \frac{p_{,\varepsilon}(\rho_r, \varepsilon)}{\rho_r} \right) = \frac{r}{\varepsilon'(T(\varepsilon))}, & \text{if } \varepsilon_r = \varepsilon_l = \varepsilon, \end{cases} \quad (\text{B.5})$$

$$\bar{\chi} = \begin{cases} \frac{\Delta p - \bar{\kappa} \Delta \rho \varepsilon}{\Delta \rho}, & \text{if } \rho_r \neq \rho_l, \\ \frac{1}{2} \left(p_{,\rho}(\rho, \varepsilon_l) - \frac{\varepsilon_l}{\rho} p_{,\varepsilon}(\rho, \varepsilon_l) + p_{,\rho}(\rho, \varepsilon_r) - \frac{\varepsilon_r}{\rho} p_{,\varepsilon}(\rho, \varepsilon_r) \right) \\ = \frac{1}{2} r(T(\varepsilon_l) - \varepsilon_l T'(\varepsilon_l) + T(\varepsilon_r) - \varepsilon_r T'(\varepsilon_r)), & \text{if } \rho_r = \rho_l = \rho. \end{cases} \quad (\text{B.6})$$

The definitions of the eigenvalues and right and left eigenvectors are easy to obtain and are omitted here.

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